

K4LED

Georgia Amateur Radio Astronomy Observatory

Observation Date: 06-23-2018, 0:00:00 to 23:59:00 UTC

Location: Lat. 34.42322 N, Lon. -84.49413 W (Jasper, GA 30143)

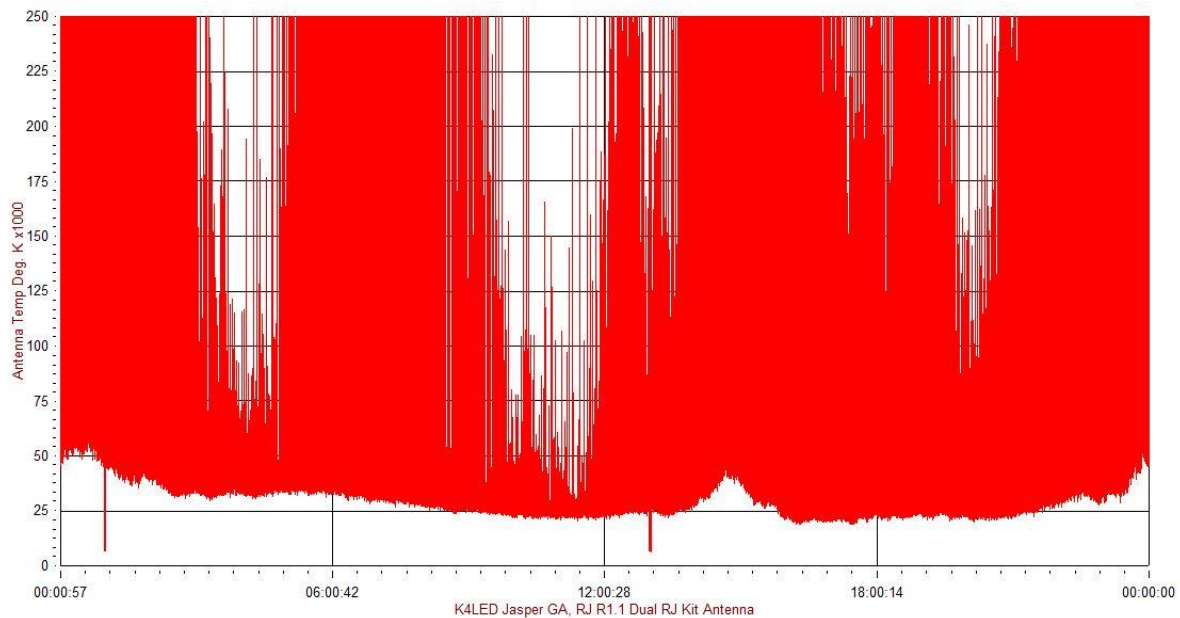
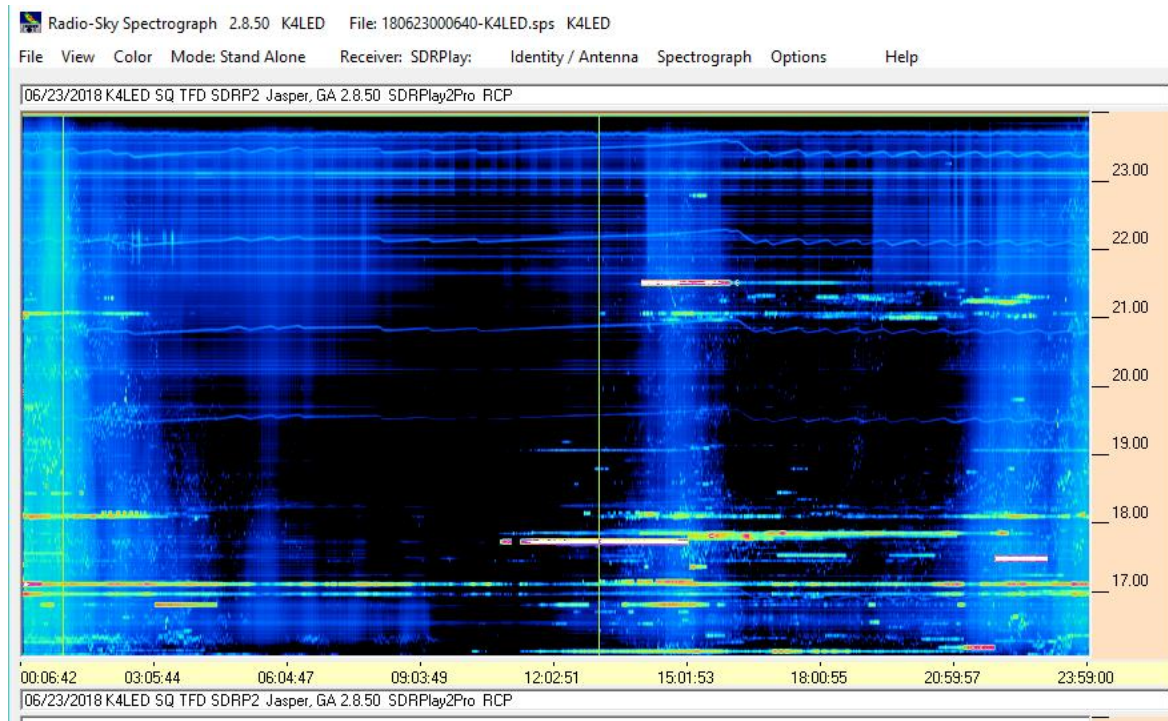
Object: TP, GB, Very Light Solar 19:18:50 (Heavy Thunderstorm)

Observer: Larry E. Dodd

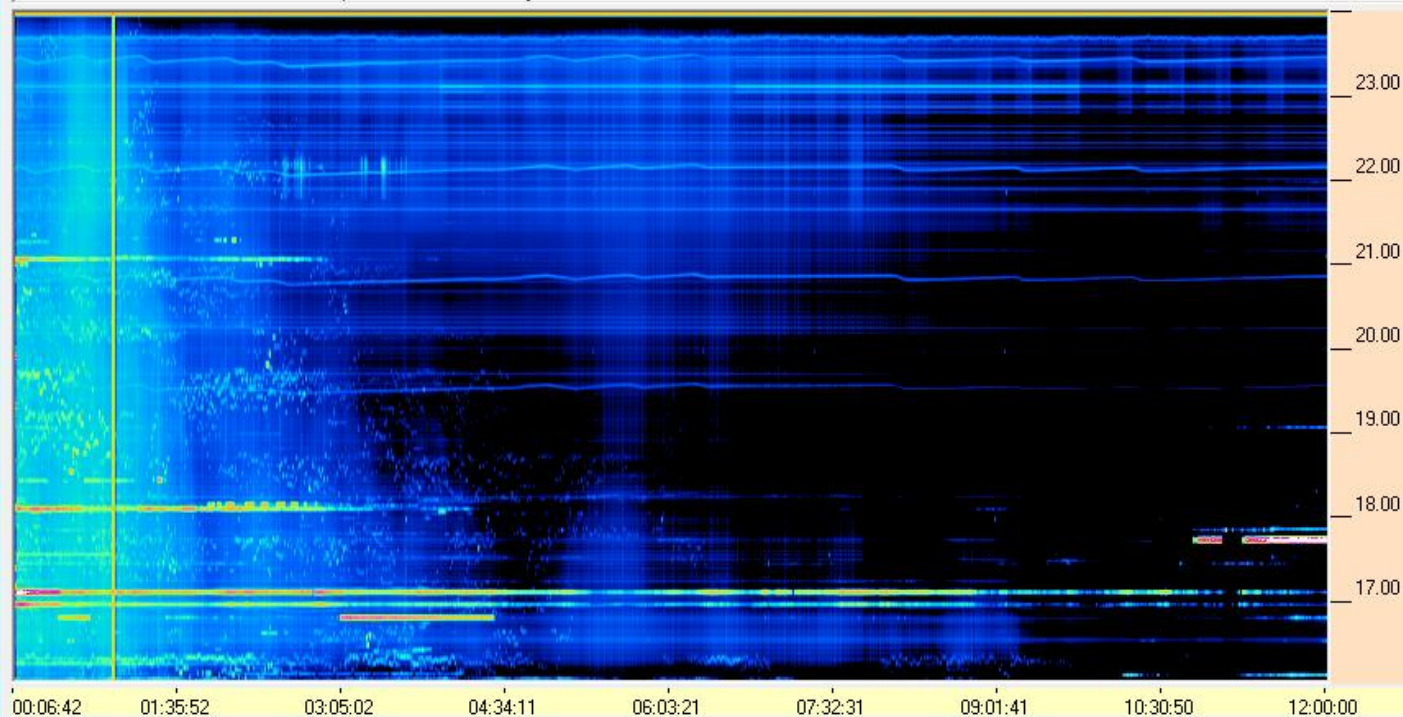
Receivers: FSX11, SDRPlay2PRO, RJ1.1

Software: SkypipeII 2.7.34, Spectrograph 2.8.50

Antennas: Radio Jove Dual Dipole Kit, Typinski Square TFD

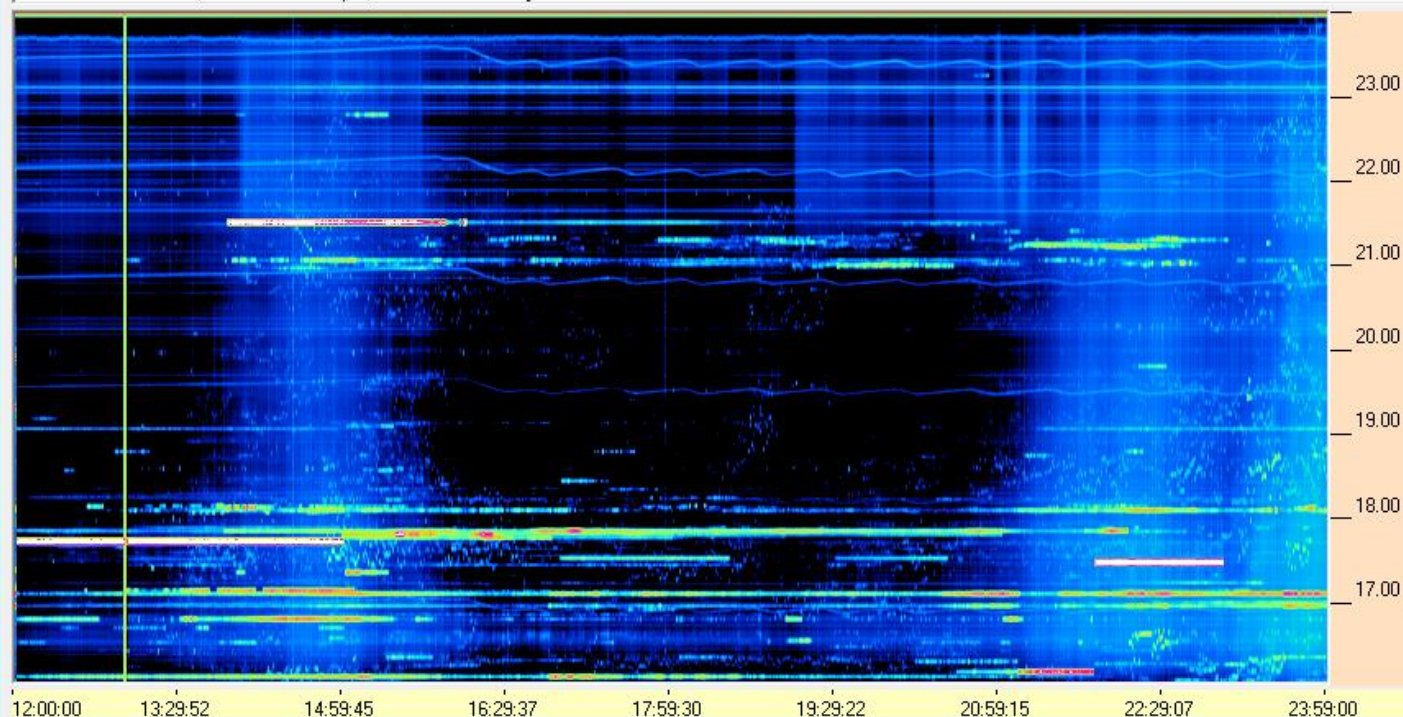


06/23/2018 K4LED SQ TFD SDRP2 Jasper, GA 2.8.50 SDRPlay2Pro RCP



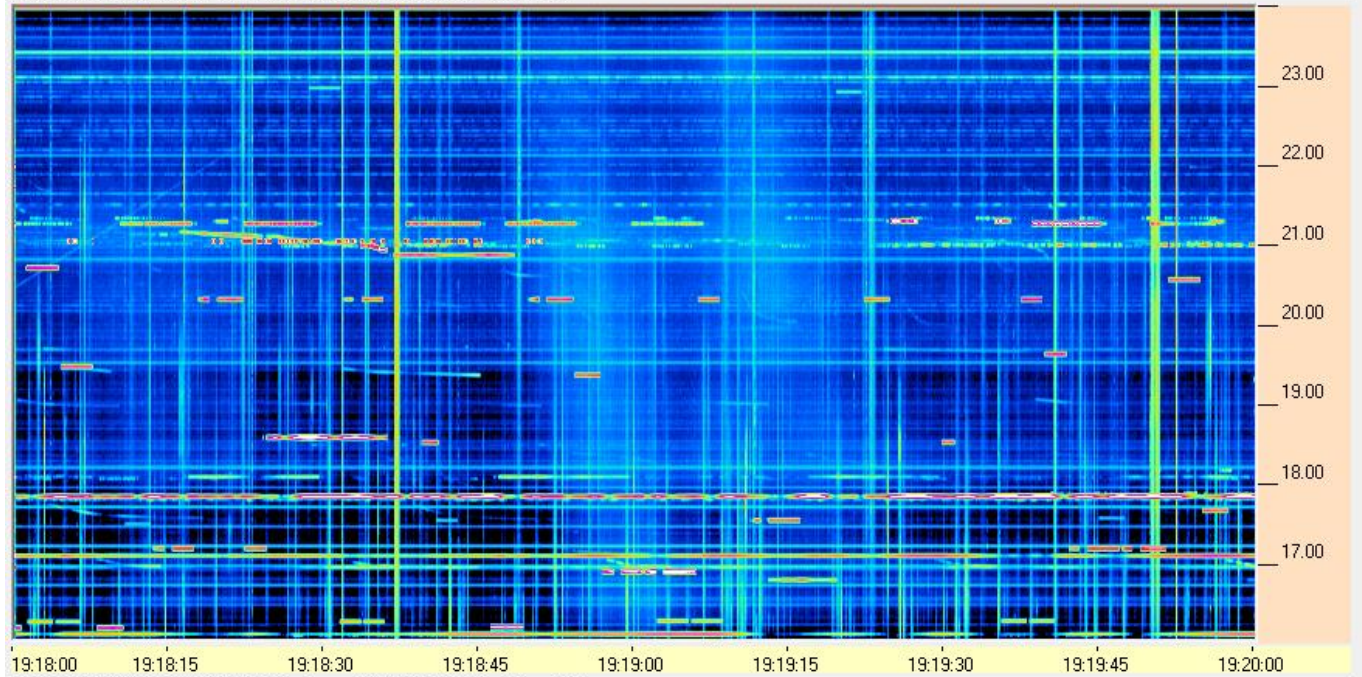
06/23/2018 K4LED SQ TFD SDRP2 Jasper, GA 2.8.50 SDRPlay2Pro RCP

06/23/2018 K4LED SQ TFD SDRP2 Jasper, GA 2.8.50 SDRPlay2Pro RCP

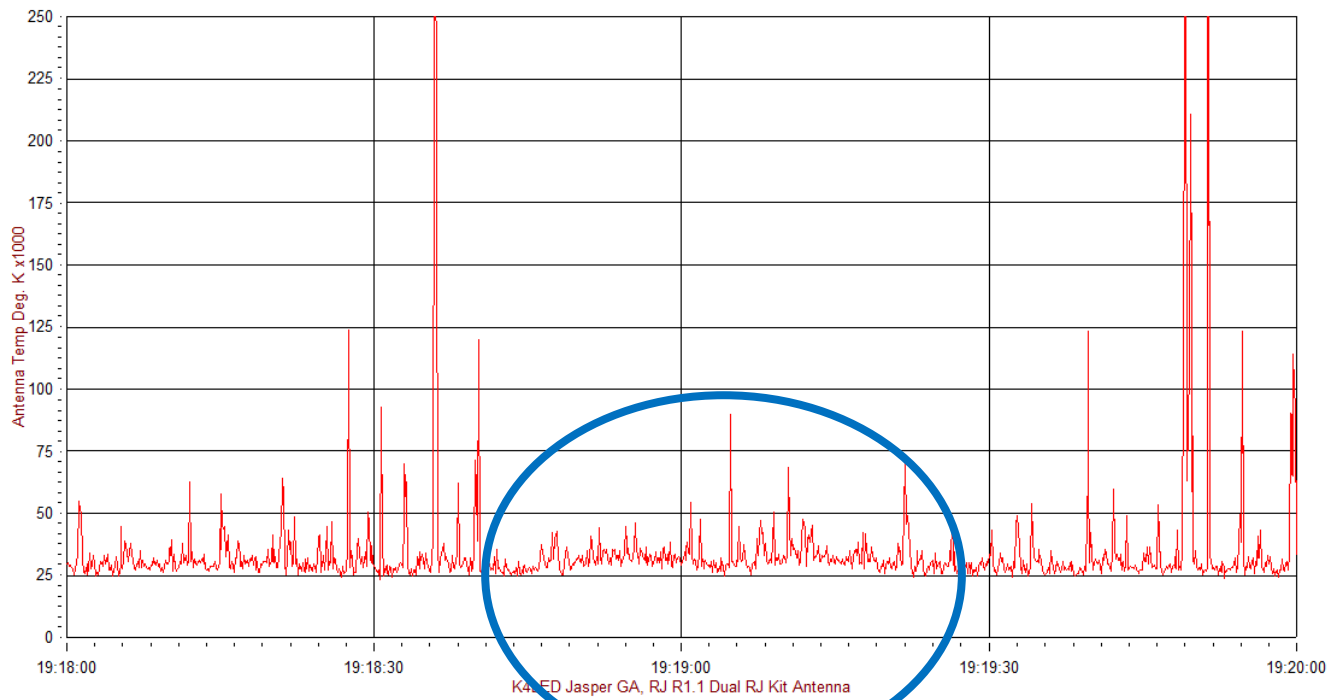


06/23/2018 K4LED SQ TFD SDRP2 Jasper, GA 2.8.50 SDRPlay2Pro RCP

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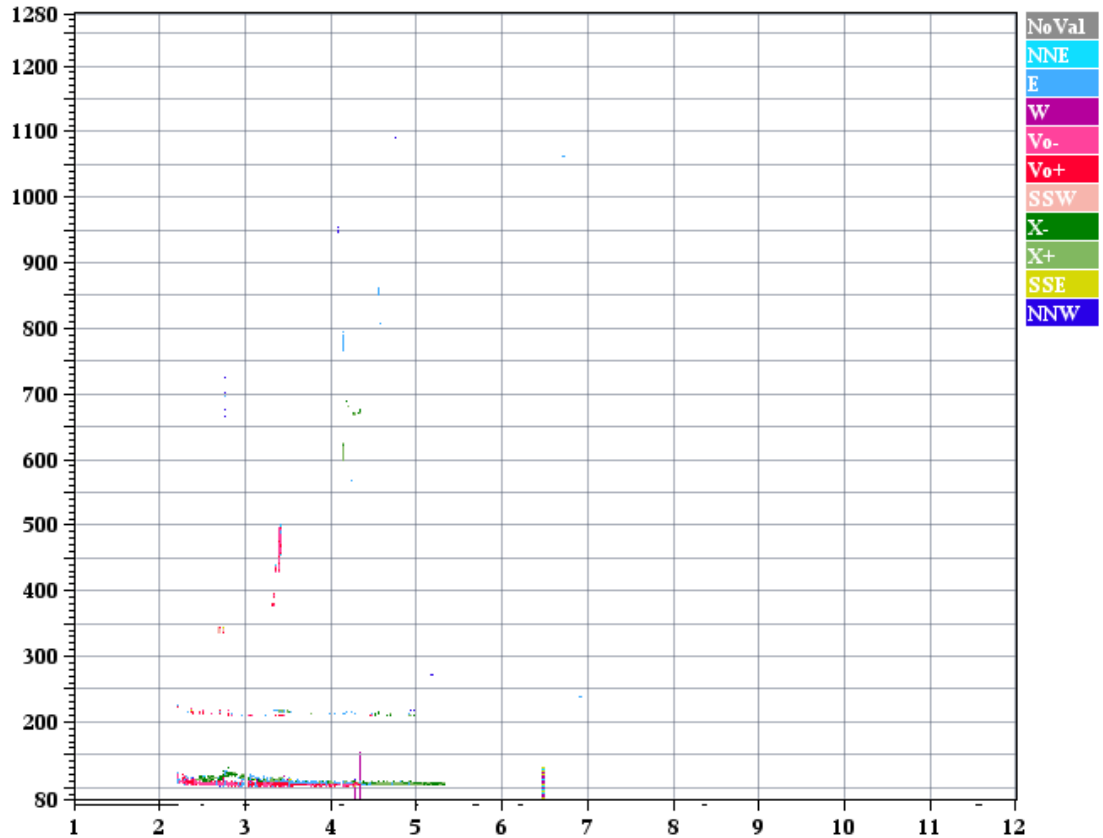


K4LED Jasper GA, RJ R1.1 Dual RJ Kit Antenna



Station YYYY DAY DDD HHMMSS P1 FFS S AXN PPS IGA PS
Eglin AFB 2018 Jun23 174 120000 RSF 1 713 100 03+ 79

foF2 N/A
foF1 N/A
foF1p N/A
foE 2.30
foEp 2.22
fxI N/A
foEs 4.45
fmin N/A
MUF(D) N/A
M(D) N/A
D N/A
h`F N/A
h`F2 N/A
h`E N/A
h`Es 102.5
hmF2 N/A
hmF1 N/A
hmE N/A
yF2 N/A
yF1 N/A
yE N/A
B0 N/A
B1 N/A
C-level N/A
Auto:



D 100 200 400 600 800 1000 1500 3000 [km]
MUF 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 [MHz]

57071860.tmp / 440fx512h 25 kHz 2.5 km / DPS-4D EG931 084 / 30.5 N 273.5 E

ShowIonogram v 1.0

References for Interpreting the Ionogram:

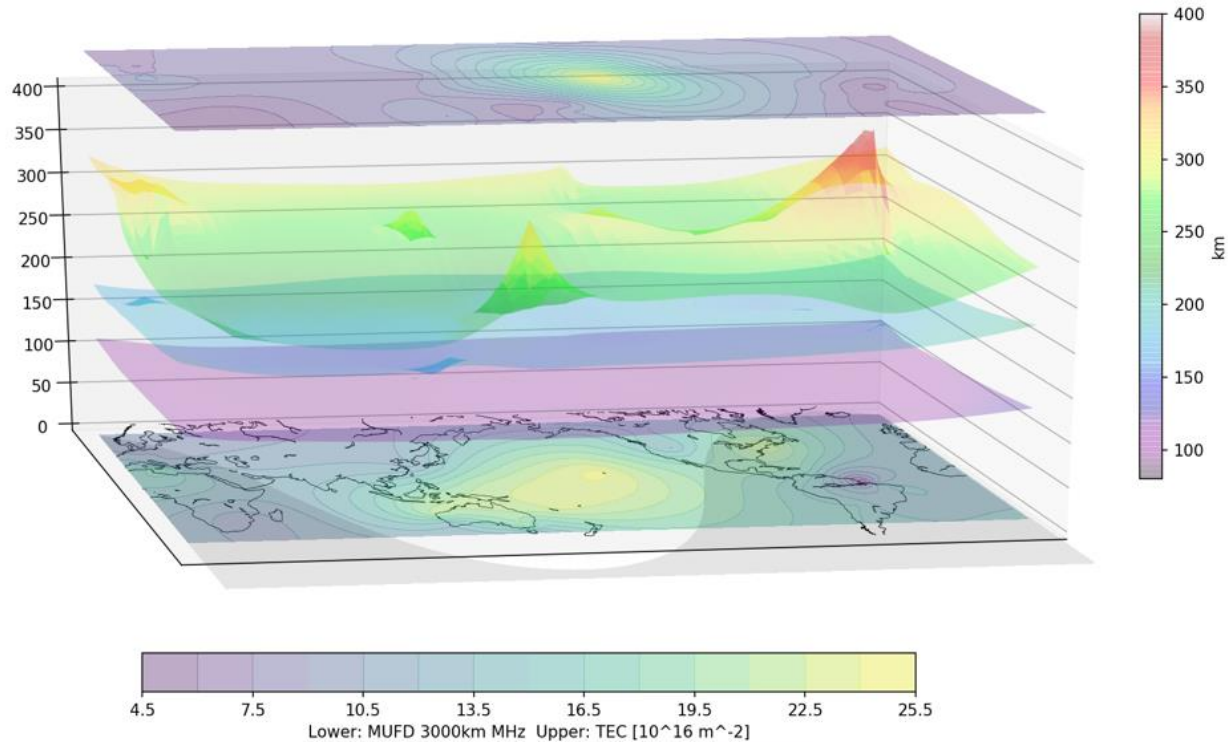
http://www.sws.bom.gov.au/IPSHosted/INAG/uag_23a/UAG_23A_indexed.pdf

<http://citeseerx.ist.psu.edu/viewdoc/download?doi=10.1.1.404.5517&rep=rep1&type=pdf>

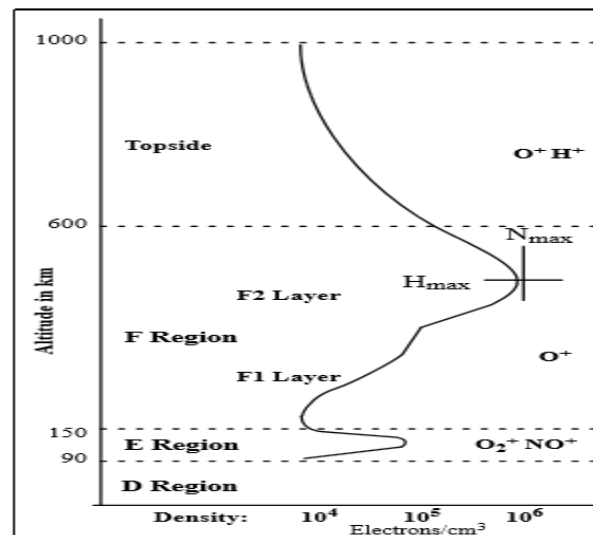
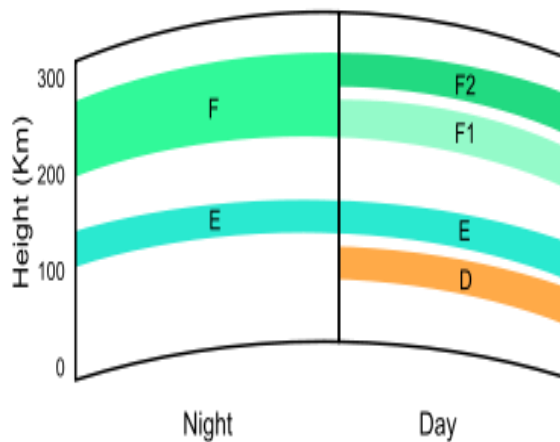
<https://www.hfunderground.com/wiki/index.php/Ionosonde>

Ionosphere Layers, MUF, Total Electron Content

MUF TEC E, F1, F2 Max Height 2018-06-24 02:00



Data Global Ionospheric Radio Observatory (GIRO) Principal Investigator Prof. B. W. Reinisch UMass Lowell



The **critical frequency** is the limiting frequency at or below which a radio wave is reflected by an ionospheric layer at vertical **incidence**. If the transmitted frequency is higher than the **plasma frequency** of the ionosphere, then the electrons cannot respond fast enough, and they are not able to re-radiate the signal. It is calculated as shown below:

$$f_{\text{critical}} = 9 \times \sqrt{N}$$

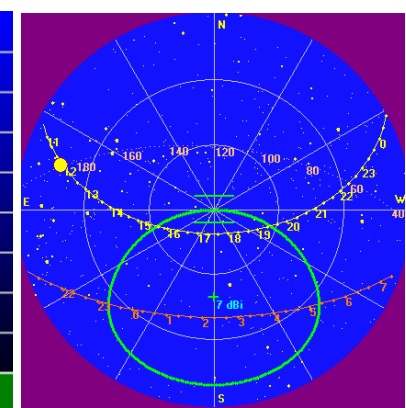
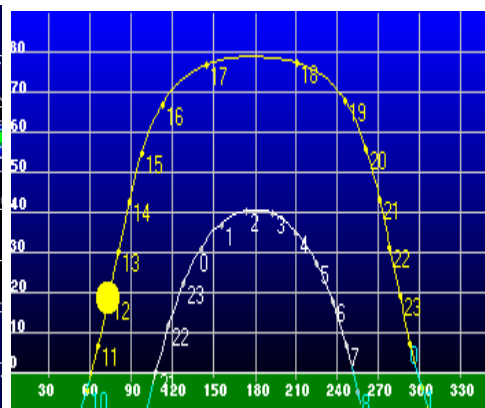
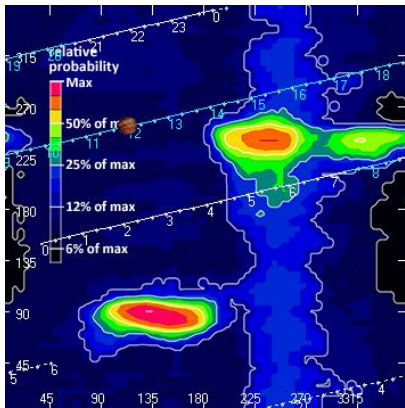
where N = electron density per m^3 and f_{critical} is in Hz.

The Maximum Usable Frequency (MUF) is defined as the upper frequency limit that can be used for transmission between two points at a specified time.

$$f_{\text{muf}} = \frac{f_{\text{critical}}}{\sin \alpha}$$

where α = **angle of attack**, the angle of the wave relative to the **horizon**, and sin is the **sine** function.

Reference: <https://en.wikipedia.org/wiki/Ionosphere>



Radio-Jupiter Pro 3.8.2 (GMT)

File Options Display Tools Window Help

Alt 02

Run

YYYY MM DD

Date 2018 06 23 cal

HH MM SS

Time 12 00 00 GMT

update

Using Fixed Time

Lat: 34:25:23.5
Long: 84:29:38.8 W

Jupiter

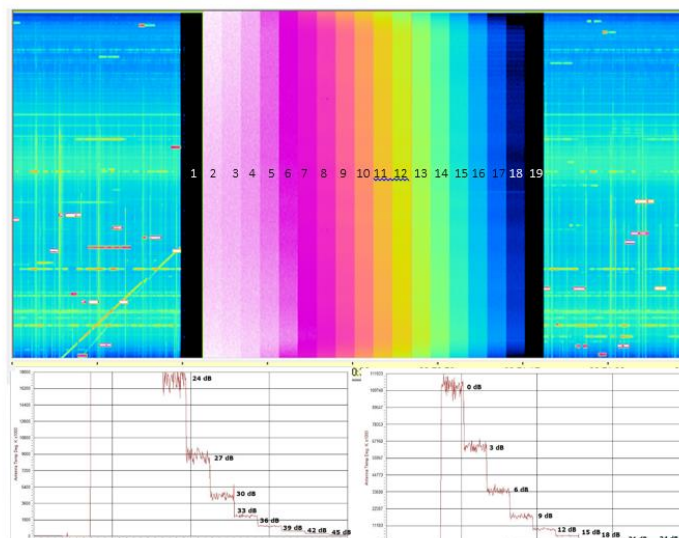
ALT: -53.4
AZ: 293.7
HA: -14:18
Rise: 21:01
Set: 07:38
RA: 14:46
DEC: -14:50
CMIH: 113.09
Io SGC: 252.64

Sun

ALT: 18.5
AZ: 073.8
HA: -05:32
Rise: 10:25
Set: 00:43
RA: 06:01
DEC: 23:26

LMST: 00:28:28

AUTOMATED CALIBRATION COLOR BAR TEST PATTERN AND DATA ACTUAL MEASURED DATA



STEP	TEMPERATURE	ATTENUATION
1	>1 kK	101 dB
2	104.0 MK	0 dB
3	66.9 MK	3 dB
4	36.3 MK	6 dB
5	19.1 MK	9 dB
6	9.6 MK	12 dB
7	4.5 MK	15 dB
8	2.3 MK	18 dB
9	1.2 MK	21 dB
10	602.0 kK	24 dB
11	296.6 kK	27 dB
12	133.0 kK	30 dB
13	92.2 kK	33 dB
14	51.6 kK	36 dB
15	32.0 kK	39 dB
16	22.6 kK	42 dB
17	17.9 kK	45 dB
18	12.0 kK	48 dB
19	>1 kK	101 dB